# Forward hadron production at collider energies and its possible application to cosmic ray physics

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# Plan

### • Motivation

high-energy hadron-hadron collisions

vs first impact of cosmic ray air shower

### Necessary ingredients for the calculation

1. Evolution eq.: rcBK - a new working paradigm

- 2. ``Initial Cond." : AAMQS DIS global fit (see talk by J.G.Milhano)
- 3. Formula: DHJ -forward hadron production
- 4. MC treatment of nuclei
- MC-DHJ/rcBK: state-of-the-art calculation work done with Fujii, Kitadono and Nara
- Future prospect towards application to CRs
- Summary

## Motivation

• The first impact of cosmic ray air showers is an **extremely high energy scattering** of (probably) a proton off a nucleus in the air.

$$E_{\text{lab}} = 10^{20} \text{eV}, \quad \sqrt{s_{pp}} = 433 \text{ TeV} \implies \sqrt{s_{pp}} = 14 \text{ TeV} \text{ (LHC)}$$



- <u>Hadronic interaction in the MC code</u> is crucial for determination of composition and correct energy estimation of the primary cosmic rays. (see talk by T. Pierog)
- We, theorists (working on high-energy scatt.), must provide up-to-date information on <u>forward hadronic cross section</u> based on the modern picture of theory.

## Aim of this talk

• To present (one of) the continuous efforts towards precise understanding of forward hadronic cross sections at high energies. (hadron production in hadron-hadron scattering)

• In particular, based on hard QCD picture. (could be irrelevant at very large rapidities, but not clear beyond which rapidity soft physics becomes dominant)

## What is necessary?



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## What is necessary?



What we really need at forward rapidity is

- 1. Evolution equation to go to higher energy (smaller x, instead of DGLAP)
- 2. Parton distribution in the target (instead of ordinary pdf's)
- 3. Formula useful at forward rapidity (instead of the collinear fact. formula)
- 4. A particular method of treating nuclei (instead of a homogeneous nucleus)

 $\begin{array}{l} x_2 \text{ becomes very small} \\ pp @ LHC (7TeV) \\ x_1, x_2 \sim 3 \times 10^{-4} \, (\eta = 0, p_t = 2 \text{GeV}) \, , \\ x_2 \sim 7 \times 10^{-7} \, (\eta = 6, p_t = 2 \text{GeV}) \\ pp @ cosmic ray (400TeV) \\ x_1, x_2 \sim 5 \times 10^{-6} \, (\eta = 0, p_t = 2 \text{GeV}) \, , \\ x_2 \sim 1 \times 10^{-8} \, (\eta = 6, p_t = 2 \text{GeV}) \end{array}$ 

## CGC appears at small x

See talks by Milhano and Blaizot



 $Q_{s}(x,A)$ : saturation momentum boundary btw saturated and NONsaturated regimes

$$Q_{\rm s} \sim 1 {\rm GeV}$$
 at  $x = 10^{-4}$  for a proton

### Going up higher energies: evolution eqs.

**Evolution wrt** x (or rapidity  $y = \ln 1/x$ ) for unintegrated gluon distribution

• **BFKL** (LO : 
$$(\alpha_s \ln 1/x)^n$$
, NLO:  $\alpha_s (\alpha_s \ln 1/x)^n$ )  

$$\frac{\partial \phi(\mathbf{x}, \mathbf{k_t})}{\partial \ln(\mathbf{x_0}/\mathbf{x})} \approx \mathcal{K} \otimes \phi(\mathbf{x}, \mathbf{k_t})$$

$$K : \text{gluon splitting } \mathbf{g} \rightarrow \text{gg}$$

$$\phi: \text{ unintegrated gluon distr.}$$

$$\frac{\partial \phi(\mathbf{x}, \mathbf{k_t})}{\partial \ln(\mathbf{x_0}/\mathbf{x})} \approx \mathcal{K} \otimes \phi(\mathbf{x}, \mathbf{k_t}) - \phi(\mathbf{x}, \mathbf{k_t})^2$$

Known up to full NLO accuracy. [Balitsky, Chirilli 2008] But for practical purposes, we use BK with running coupling  $\rightarrow$  "rcBK" [Balitsky, Chirilli 2008]

[Balitsky, Gardi et al., Kovchegov-Weigert]

$$K^{\rm run}(\mathbf{r}, \mathbf{r_1}, \mathbf{r_2}) = \frac{N_c \,\alpha_s(r^2)}{2\pi^2} \left[ \frac{r^2}{r_1^2 \, r_2^2} + \frac{1}{r_1^2} \left( \frac{\alpha_s(r_1^2)}{\alpha_s(r_2^2)} - 1 \right) + \frac{1}{r_2^2} \left( \frac{\alpha_s(r_2^2)}{\alpha_s(r_1^2)} - 1 \right) \right]$$

## Fit to HERA data: AAMQS<sub>2011</sub>

Initial Conditions : modified GBW/MV models  $x_0 = 0.00893$  or 0.008 $\mathcal{N}^{\text{GBW}}(r, x = x_0) = 1 - \exp\left[-\frac{(r^2 Q_{s0}^2)^{\gamma}}{4}\right],$  $(\gamma = 1 : \text{ordinary GBW})$  $\mathcal{N}^{\text{MV}}(r, x = x_0) = 1 - \exp\left[-\frac{(r^2 Q_{s0}^2)^{\gamma}}{4} \ln\left(\frac{1}{rA} + e\right)\right]$  $(\gamma = 1 : \text{ordinary MV})$ IR regularization for 1-loop running coupling  $\alpha_{s,n_f}(r^2) = \frac{4\pi}{\beta_{0,n_f} \ln\left(\frac{4C^2}{r^2 A^2}\right)}$ freeze the coupling at  $\alpha_{\rm s}^{\rm fr}=0.7$ Fit with only light quarks Fit including heavy quarks Q<sup>2</sup>=0.85 GeV<sup>2</sup> Q<sup>2</sup>=0.85 GeV<sup>2</sup> Data Data Q<sup>2</sup>=2.0 GeV<sup>2</sup>  $Q^2 = 2.0 \text{ GeV}^2$ ← Modified GBW  $\sigma_r$ Q<sup>2</sup>=8.5 GeV<sup>2</sup>  $Q^2=4.5 \text{ GeV}^2$  $Q^2 = 4.5 \text{ GeV}$ Q<sup>2</sup>=8.5 GeV<sup>2</sup> (Left) γ=0.971  $\sigma_{r_0}$  $\sigma_{r_0}$  $Qs_0^2 = 0.241$ Q<sup>2</sup>=12.0 GeV Q<sup>2</sup>=12.0 GeV Q2=10.0 GeV2 Q<sup>2</sup>=10.0 GeV<sup>2</sup>  $\sigma_{\rm r_{0.5}}$  $\sigma_{\rm r_{0.5}}$ (Right)  $\gamma$ =0.959 O<sup>2</sup>=28.0 GeV Q<sup>2</sup>=15.0 Ge\ Q<sup>2</sup>=28.0 GeV Q<sup>2</sup>=15.0 GeV  $Qs_0^2 = 0.240$  $\sigma_{r_0}$  $\sigma_{r_0}$ 

Q<sup>2</sup>=35 GeV<sup>2</sup>

10<sup>-4</sup>

 $\sigma_{r_{0,r}}$ 

10-

Q<sup>2</sup>=45 GeV<sup>2</sup>

10<sup>-4</sup>

 $m_{\rm c}$ =1.27GeV,  $m_{\rm b}$ =4.2GeV

x 10<sup>-2</sup>

10<sup>-3</sup>

Q<sup>2</sup>=35 GeV<sup>2</sup>

 $10^{-4}$ 

10<sup>-3</sup>

10-

m<sub>uds</sub>=140MeV

 $\sigma_{r_{0,r}}$ 

10

Q<sup>2</sup>=45 GeV<sup>2</sup>

10<sup>-4</sup>

10<sup>-3</sup>

x 10<sup>-2</sup>

There are two more parameters ( $C, \sigma_0$ )

## Hadron collisions (pp/pA): two formulae

#### *k*<sub>t</sub> factorization

$$\frac{d\sigma^{A+B\to g}}{dyd^2p_Td^2X} \sim K \frac{\alpha_s}{p_T^2} \phi_{\mathsf{A}}(k_1, x_1, b) \otimes \phi_{\mathsf{B}}(k_2, x_2, X - b)$$

- proved for pp, pA at LO
- good when both A and B are saturated (mid rapidity at very high energy)
- used in various calculations e.g. multiplicity distribution, etc



DHJ formalism [Dumitru-Hayashigaki-Jalilian--Marian 2006]

$$\frac{dN}{dy_h d^2 p_T} = \frac{K}{(2\pi)^2} \sum_{ijk} \int_{x_F}^1 \frac{dz}{z^2} x_1 f_{i/p}(x_1, p_T^2) \,\widetilde{\mathcal{N}}_j(\frac{p_T}{z}, x_2) \, D_{h/k}(z, p_T^2)$$

- "Large-x / small-x" reactions: valid at forward rapidity
   x<sub>1</sub>~1, x<sub>2</sub> <<1</li>
- $-f_{i/p}(x)$  : pdf for valence (large x) partons in the projectile
- $D_{h/k}(z)$  : frag. func. for outgoing hadron h from a parton k
- N : un-integrated gluon distribution in the target



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## How to treat nuclei?

MC modeling for a nucleus:

• The simplest will be a homogeneous disk no impact parameter dependence an additional parameter  $Q_{s0A}^2$  needed

may use a simple parametrization by KLN, or numerical solution to rcBK

 Random nucleons w/ Woods-Saxon dist. fluctuating density ⇒ <u>b-dependence</u> Q<sup>2</sup><sub>s0A</sub> = Q<sup>2</sup><sub>s0p</sub> × N w/o additional parameter Drescher-Nara



Apply quantum evolution locally at different transverse bins



# MC-DHJ/rcBK

[Fujii,KI,Kitadono,Nara, arXiv:1107.1333, more to come soon]

To reduce ambiguity

- construct a nucleus by randomly placing nucleons
- use AAMQS parameters for proton IC optimized for DIS at small-x
- quantum evolution is performed "locally" in b space with rcBK

(to avoid IR div. in b-dep BK)



## **MC-DHJ/rcBK : results**



modified MV model ( $\gamma = 1.118$ )

#### "running coupling" version of MV

**model** [Iancu-KI-Triantafylopoulos] : to be consistent with rcBK evolution

- reproduce the data nicely
- $\bullet$  AAMQS set h and rcMV for  $\mathcal{N}(r,y)$
- $Q_{s0A}^2$  fixed by MC; no additional parameter

## **MC-DHJ/rcBK : results**

dN/dŋd<sup>2</sup>p<sub>T</sub> [GeV<sup>-2</sup>



## MC-DHJ/rcBK extrapolated to LHC



• Hadron productions  $(\pi^0, K^0 \text{ and } n)$  at  $\eta = 8.5$  at 7 TeV (LHCf) is being studied in this framework

Very forward region could be dominated by soft interaction, but still necessary to understand how much hard contribution exists.

### **Future prospect towards application to CRs**

- Separation between soft and hard is not clear (model dependent)
- In several hadron interaction models (e.g. SIBYLL), IR cutoff for the hard contribution is energy dependent (very similar to Qs(s))

SIBYLL2.0  $p_{\perp}^{cutoff} = p_{\perp}^{0} + \Lambda \exp\left\{c\sqrt{\ln(s/\text{GeV}^{2})}\right\}$ with  $p_{\perp}^{0} = 1 \text{ GeV}, \Lambda = 0.065 \text{ GeV}$  and c = 0.9

 $\rightarrow$  miss particle production in the semi-hard region!!

• CGC provides particle production in semi-hard region  $\Lambda_{\rm QCD} < k_t < Q_{\rm s}$  that expands with increasing energy

 $\rightarrow$  filling the gap btw soft and hard

Calculations with CGC could help to "recover" the semi-hard contributions.

## Summary

- Theoretical description of high-energy hadron scattering based on CGC is now (almost) established up to leading log accuracy with running coupling corrections. → rcBK paradigm
- In particular, phenomenological analysis with rcBK has been making a progress enough to be compared with experimental data. → HERA DIS at small-x, RHIC dAu at forward rapidity
- This approach can be, in principle, applied to higher energy collision, thus hopefully to the first impact of Cosmic Rays in the air.