

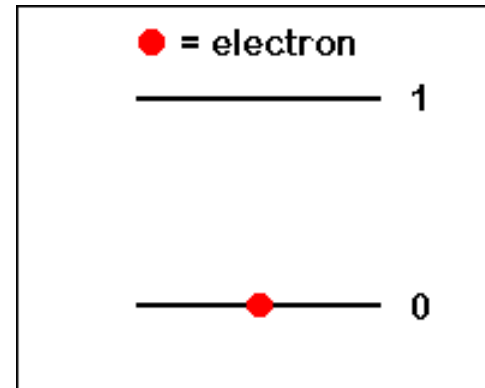
# 2. Laser physics - basics

Spontaneous and stimulated processes

Einstein A and B coefficients

Rate equation analysis

Gain saturation



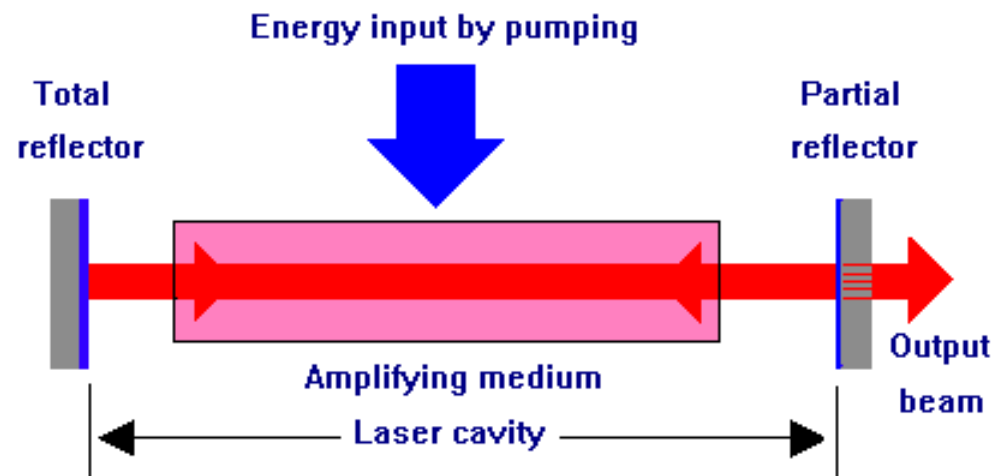
# What is a laser?

LASER: **L**ight **A**mplification by **S**timulated **E**mission of **R**adiation

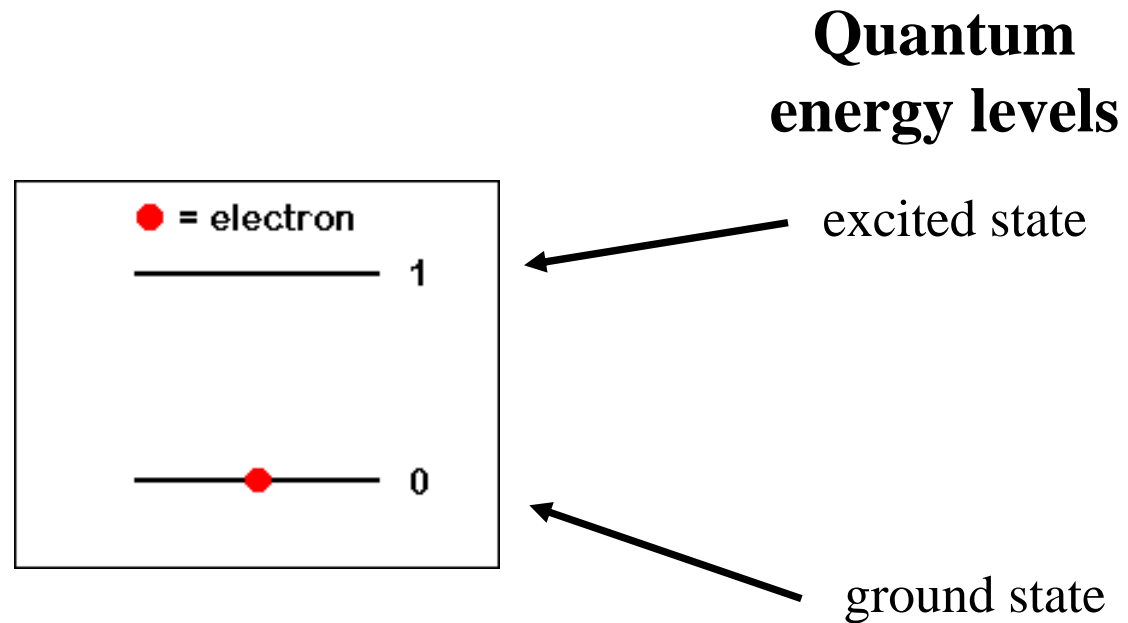
↙ "light" could mean anything from microwaves to x-rays

Essential elements:

1. A laser medium - a collection of atoms, molecules, etc.
2. A pumping process - puts energy into the laser medium
3. Optical feedback - provides a mechanism for the light to interact (possibly many times) with the laser medium



# The two-level atom



Absorption: promotes an electron from the ground to the excited state  
Emission: drops the electron back to the ground state

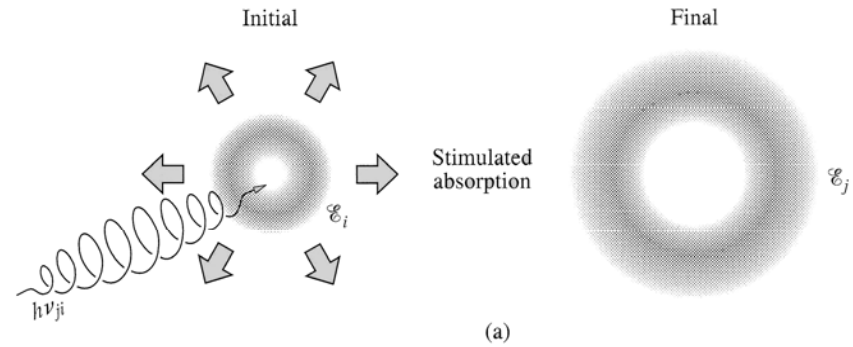
"**spontaneous emission**" - the decay of an excited state to the ground state with the corresponding emission of a photon

Conservation of energy:  $E_{\text{excited}} - E_{\text{ground}} = E_{\text{photon}}$

# Three things can occur

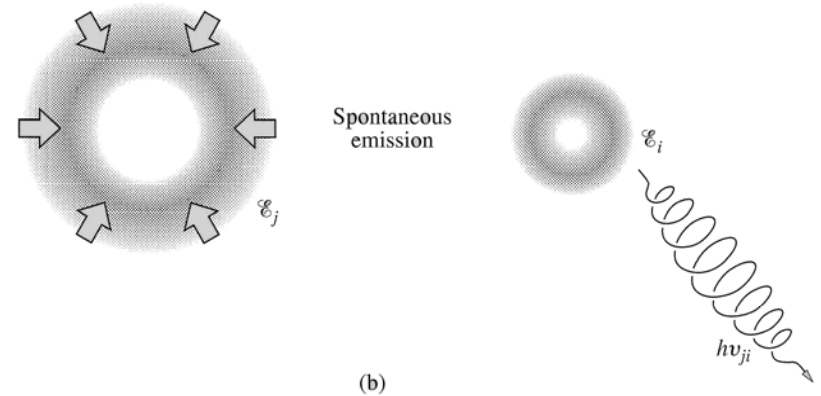
## Absorption

- Promotes molecule to a higher energy state
- Decreases the number of photons



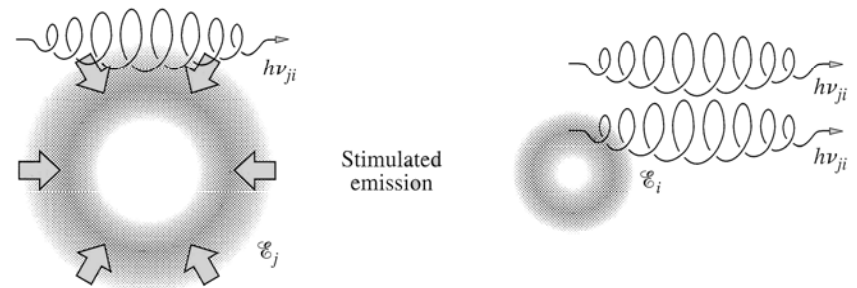
## Spontaneous Emission

- Molecule drops from a high energy state to a lower state
- Increases the number of photons
- This is the only one that does NOT require a photon in the initial state



## Stimulated Emission

- Molecule drops from a high energy state to a lower state
- The presence of one photon stimulates the emission of a second one



# Relaxation of the two-level atom

An atom in the excited state can relax to the ground state by:

- spontaneous emission: rate is  $\gamma_{\text{rad}}$
- any of a variety of non-radiative pathways: rate =  $\gamma_{\text{nr}}$

All of these processes are single-atom processes;  
each atom acts independently of all the others.

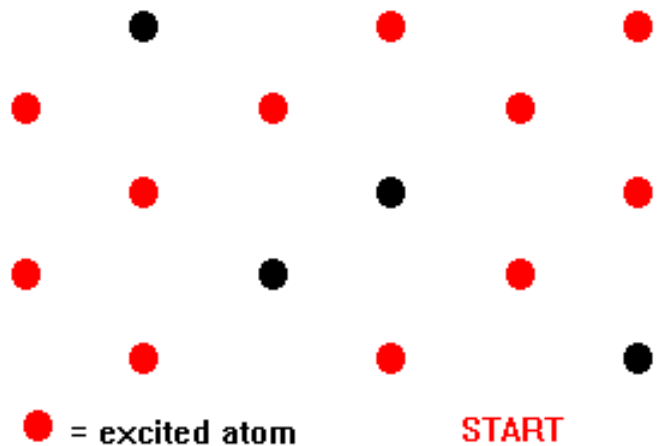
Thus, evolution of the excited state population only depends on the number of atoms in the excited state:

$$\frac{dN_e}{dt} = -\gamma_{\text{rad}}N_e - \gamma_{\text{nr}}N_e = -\gamma_{10}N_e$$

$\gamma_{10}$  = total spontaneous relaxation rate from state 1 to state 0

# A collection of two-level atoms

"Stimulated transitions" - a **collective process** involving many two-level atoms



stimulated **absorption**: light induces a transition from 0 to 1

stimulated **emission**: light induces a transition from 1 to 0

↙  
In the emission process, the emitted photon is **identical** to the photon that caused the emission!

Stimulated transitions: likelihood depends on the number of photons around

# How did it all begin?

Rayleigh-Jeans law (circa 1900):

$$\text{energy density of a radiation field } u(\nu) = 8\pi\nu^2 kT/c^3$$

Note: the units of this expression are correct. Strictly speaking,  $u(\nu)$  is an energy density per unit bandwidth, such that the integral  $\int u(\nu) d\nu$  gives an answer with units of energy/volume.

Total energy radiated from a black body:  $\int u(\nu) d\nu = \infty$

uh-oh... the "ultraviolet catastrophe"

Solution: quantum mechanics

## Time-dependent perturbation theory

As a result of a perturbation  $h(t)$ , a system in quantum state 1 makes a transition to quantum state 2 with probability given by:

$$P_{1 \rightarrow 2} = \frac{1}{\hbar^2} \left| \int_{-\infty}^t e^{iE_{21}t'/\hbar} h(t') dt' \right|^2$$

Notation:

$$\omega_{21} = \frac{E_{21}}{\hbar}$$

# Harmonic perturbation

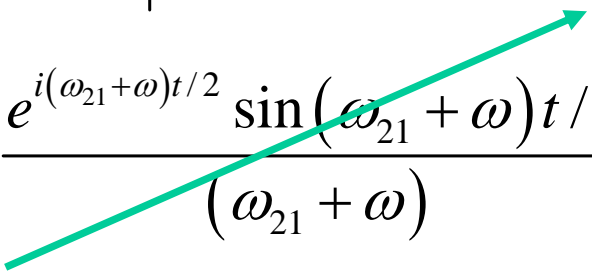
Key example: suppose we subject a two-level system, initially in state 1, to a harmonic perturbation, of the form:

$$h(t) = \begin{cases} 0 & t < 0 \\ 2A_0 \sin \omega t & t > 0 \end{cases} \quad \text{(and suppose that the frequency of the perturbation, } \omega, \text{ is close to } \omega_{21}\text{)}$$

Transition probability to state 2 is:

$$P_{1 \rightarrow 2} = \frac{A_0^2}{\hbar^2} \left| \int_0^t e^{i\omega_{21}t'} (e^{i\omega t'} - e^{-i\omega t'}) dt' \right|^2$$

$$= \frac{4A_0^2}{\hbar^2} \left| \left( \frac{e^{i(\omega_{21}-\omega)t/2} \sin(\omega_{21}-\omega)t/2}{(\omega_{21}-\omega)} \right) - \left( \frac{e^{i(\omega_{21}+\omega)t/2} \sin(\omega_{21}+\omega)t/2}{(\omega_{21}+\omega)} \right) \right|^2$$

RWA 

$$\approx \frac{4A_0^2}{\hbar^2} \cdot \frac{\sin^2 [(\omega_{21}-\omega)t/2]}{(\omega_{21}-\omega)^2}$$

Note that  $P_{1 \rightarrow 2} = P_{2 \rightarrow 1}$

Absorption and stimulated emission are **equally likely!**



# Einstein A and B coefficients

Consider a radiation field and a collection of two-level systems, in thermal equilibrium with each other.

stimulated emission probability: proportional to the number of atoms in upper state  $N_2$ , and also to the number of photons

spontaneous emission probability: proportional to  $N_2$ , but does not depend on the photon density!

$$W_{2 \rightarrow 1} = A \cdot N_2 + B \cdot u(\nu) N_2$$

Note: this is the same as  $\gamma_{\text{rad}}$

stimulated absorption probability: proportional to the number of atoms in lower state  $N_1$ , and also to the number of photons

spontaneous absorption: there is no such thing

$$W_{1 \rightarrow 2} = B \cdot u(\nu) N_1$$

Quantum mechanics says that these two coefficients must be equal!

But: in thermal equilibrium, the upward and downward transition rates must balance:

$$W_{1 \rightarrow 2} = W_{2 \rightarrow 1}$$

# Einstein A and B coefficients

Equate these two rates:

$$\frac{N_1}{N_2} = \frac{A + B \cdot u(\nu)}{B \cdot u(\nu)}$$

But Boltzmann's Law tells us that  
(in equilibrium)

$$\frac{N_1}{N_2} = e^{(E_2 - E_1)/kT}$$

Recognizing that  $E_2 - E_1 = h\nu$ , we solve for  $u(\nu)$ :

$$u(\nu) = \frac{A/B}{e^{h\nu/kT} - 1}$$

This must correspond to the Rayleigh-Jeans result in the classical limit ( $h \rightarrow 0$ ), which implies:

$$\frac{A}{B} = \frac{8\pi h\nu^3}{c^3}$$

Also has units of energy density per unit bandwidth

Since  $A = \gamma_{\text{rad}}$ , we can now solve for B also:

$$B = \frac{\gamma_{\text{rad}} c^3}{8\pi h\nu^3}$$

# Transition rates

Our expression for the downward transition rate is now:

$$\begin{aligned}W_{2 \rightarrow 1} &= A \cdot N_2 + B \cdot u(\nu) N_2 \\ &= A \cdot N_2 \left( 1 + \frac{B}{A} u(\nu) \right)\end{aligned}$$

Bose-Einstein  
distribution



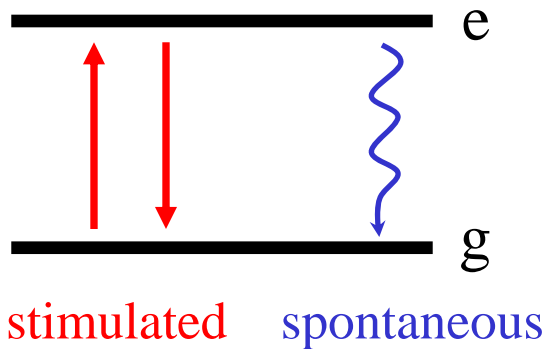
But since  $u(\nu) = \frac{A/B}{e^{h\nu/kT} - 1}$  we therefore have  $W_{2 \rightarrow 1} = A \cdot N_2 \left( 1 + \frac{1}{e^{h\nu/kT} - 1} \right)$

In other words,  $W_{21}$  is proportional to: 1 + the number of photons.

It is easy to see that the upward transition rate,  $W_{12}$ , is proportional to the number of photons:

$$W_{1 \rightarrow 2} = B \cdot u(\nu) N_1 = AN_1 \left( \frac{1}{e^{h\nu/kT} - 1} \right)$$

# Rate equation analysis



spontaneous emission:  
proportional to initial state population

stimulated transitions:  
proportional to initial state population  
proportional to photon density  $n_p$   
the same for upwards, downwards transitions

$$\frac{dN_e}{dt} = -\gamma_{eg} N_e - Kn_p N_e + Kn_p N_g = -\frac{dN_g}{dt}$$

$$\frac{dN_e}{dt} = -\gamma_{eg} N_e + Kn_p (N_g - N_e)$$

Note: the constant  $K$  is simply given by  $h\nu \cdot B$ , where  $B$  is the Einstein B coefficient

## Rate equation analysis, continued

$$\frac{dN_e}{dt} = -\gamma_{eg} N_e + Kn_p (N_g - N_e)$$

emitted photons go  
in all directions

emitted photons go only into the  
direction of the incident light

So, photon number varies according to:

$$\frac{dn_p}{dt} = -Kn_p (N_g - N_e) = -Kn_p \cdot \Delta N$$

—————> Number of photons grows **exponentially** if  $N_e > N_g$

**A LASER!**

## Rate equation analysis, part 3

In thermal equilibrium:

$$N_e = e^{-\Delta E/kT} \cdot N_g < N_g$$

→ Population inversion is impossible in equilibrium.

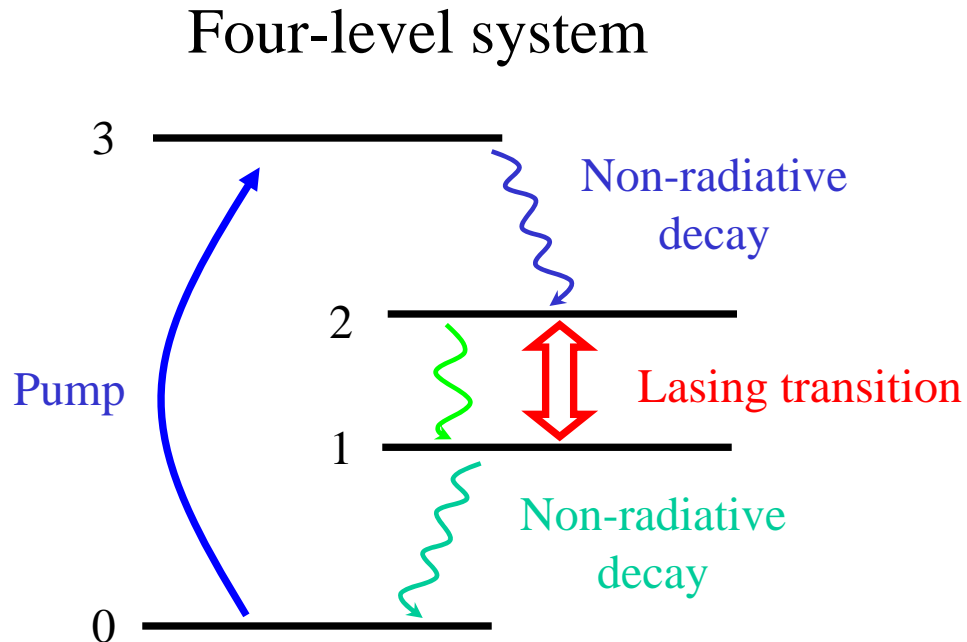
In a steady-state situation:

$$\frac{dN_e}{dt} = -\gamma_{eg} N_e + Kn_P (N_g - N_e) = 0$$

$$N_e = \frac{Kn_P}{\gamma_{eg} + Kn_P} N_g < N_g$$

→ Population inversion is impossible in steady-state.

# So how do you make a laser?



Steps 1 and 2:

Combine to give an *effective* pumping rate for level 2:  $R_p$

Step 3:

stimulated transitions due to  $n_p$   
spontaneous decay rate:  $\gamma_{21}$

Step 4:

spontaneous decay rate:  $\gamma_{10}$

$$\frac{dN_2}{dt} = R_p - \gamma_{21}N_2 + Kn_p(N_1 - N_2)$$

$$\frac{dN_1}{dt} = \gamma_{21}N_2 - Kn_p(N_1 - N_2) - \gamma_{10}N_1$$

$$\frac{dN_0}{dt} = \gamma_{10}N_1 - R_p$$

# The four-level model

Steady-state solution:

$$\Delta N = N_1 - N_2 = \frac{R_P}{\gamma_{10}} \left[ \frac{\gamma_{21} - \gamma_{10}}{\gamma_{21} + Kn_P} \right]$$

Population inversion (i.e.,  $\Delta N < 0$ ) is assured if  $\gamma_{10} > \gamma_{21}$

(even if  $n_p = 0$ , and even if  $R_p$  is small)

A necessary condition for lasing

Other necessary conditions:

- a resonant cavity - provides feedback
- net gain per round trip  $>$  net loss per round trip - “threshold”



# Saturation in the four-level atom

$$\Delta N = \frac{R_P}{\gamma_{10}} \left[ \frac{\gamma_{21} - \gamma_{10}}{\gamma_{21} + Kn_p} \right] = \frac{\gamma_{21} - \gamma_{10}}{\gamma_{21}\gamma_{10}} \cdot R_P \left[ \frac{1}{1 + W_{sig} \tau_{21}} \right]$$

- population inversion when  $\tau_2 > \tau_1$
- small signal inversion is proportional to the pump rate
- inversion level drops when  $W_{sig} > \gamma_{21}$
- the characteristic intensity for this effect is independent of pump rate  $R_p$



→ "gain saturation"

Note:  $W_{sig}$  is proportional to  $n_p$  and therefore to the intensity of the light in the medium. Thus, we can define a **saturation intensity**  $I_{sat}$  such that:

$$W_{sig} \tau_{21} = I / I_{sat}$$

