

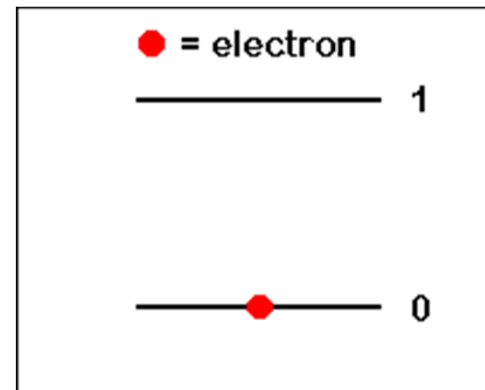
2. Laser physics - basics

Spontaneous and stimulated processes

Einstein A and B coefficients

Rate equation analysis

Gain saturation



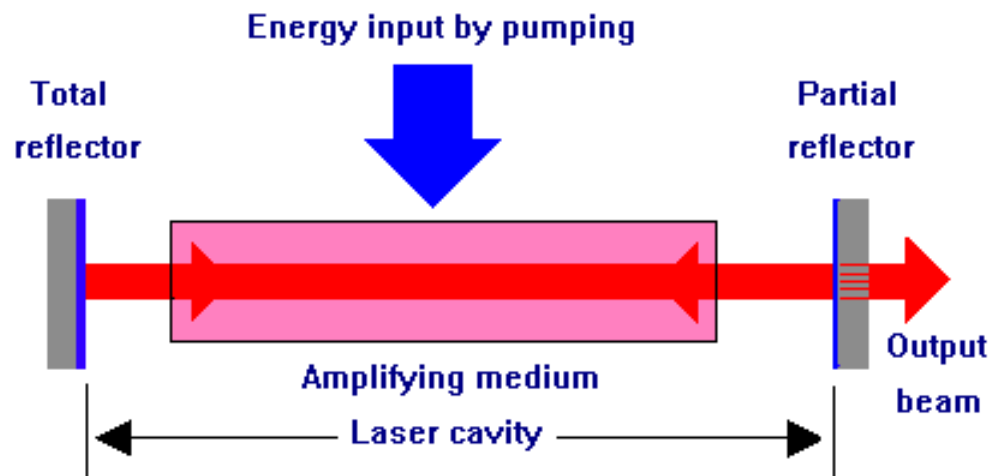
What is a laser?

LASER: **L**ight **A**mplification by **S**timulated **E**mission of **R**adiation

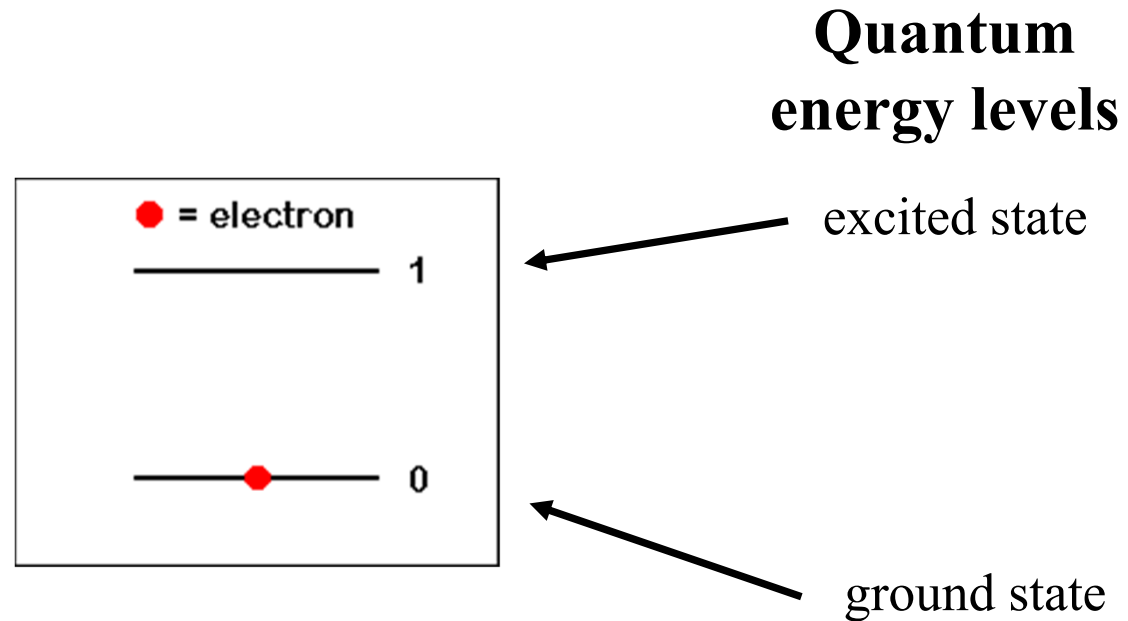
↙ "light" could mean anything from microwaves to x-rays

Essential elements:

1. A laser medium - a collection of atoms, molecules, etc.
2. A pumping process - puts energy into the laser medium
3. Optical feedback - provides a mechanism for the light to interact (possibly many times) with the laser medium



The two-level atom



Absorption: promotes an electron from the ground to the excited state

Emission: drops the electron back to the ground state

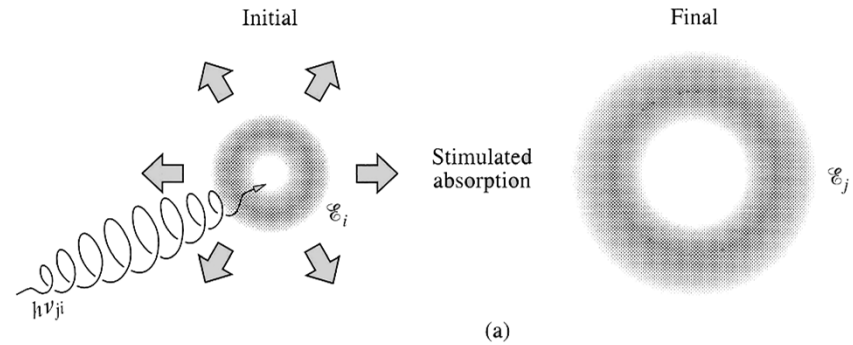
"**spontaneous emission**" - the decay of an excited state to the ground state with the corresponding emission of a photon

Conservation of energy: $E_{\text{excited}} - E_{\text{ground}} = E_{\text{photon}}$

Three things can occur

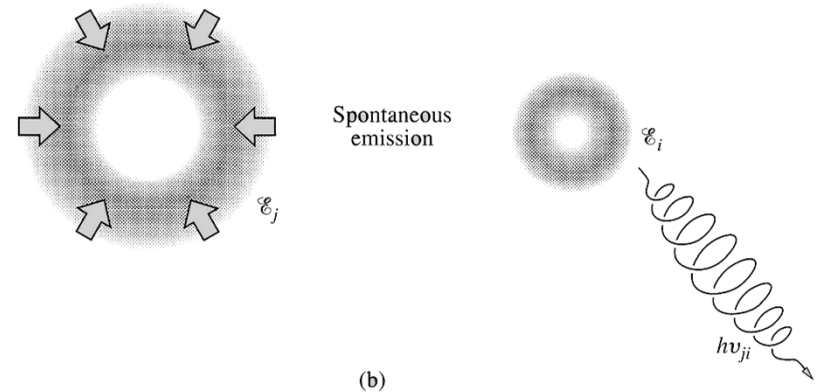
Absorption

- Promotes molecule to a higher energy state
- Decreases the number of photons



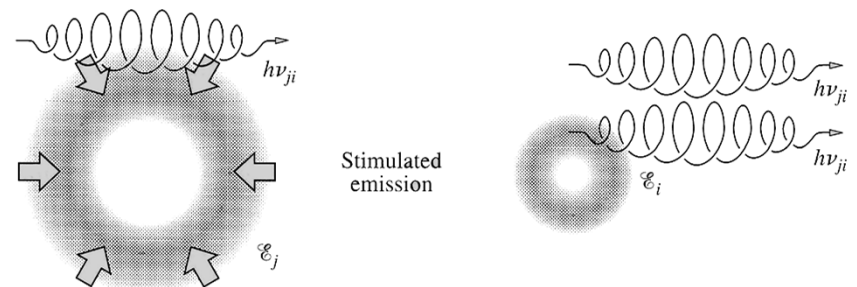
Spontaneous Emission

- Molecule drops from a high energy state to a lower state
- Increases the number of photons
- This is the only one that does NOT require a photon in the initial state



Stimulated Emission

- Molecule drops from a high energy state to a lower state
- The presence of one photon stimulates the emission of a second one



Relaxation of the two-level atom

An atom in the excited state can relax to the ground state by:

- spontaneous emission: rate is γ_{rad}
- any of a variety of non-radiative pathways: rate = γ_{nr}

All of these processes are single-atom processes;
each atom acts independently of all the others.

Thus, evolution of the excited state population only depends on the number of atoms in the excited state:

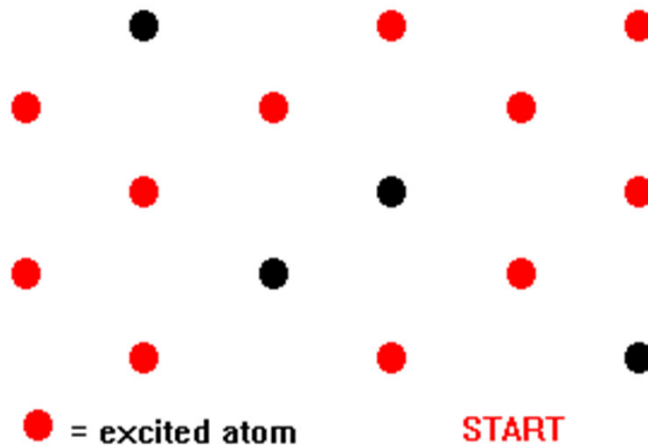
$$\frac{dN_e}{dt} = -\gamma_{\text{rad}}N_e - \gamma_{\text{nr}}N_e = -\gamma_{10}N_e$$

γ_{10} = total spontaneous relaxation rate from state 1 to state 0

For today's analysis, we ignore γ_{nr} .

A collection of two-level atoms

"Stimulated transitions" - a **collective process** involving many two-level atoms



stimulated **absorption**: light induces a transition from 0 to 1

stimulated **emission**: light induces a transition from 1 to 0

→ In the emission process, the emitted photon is **identical** to the photon that caused the emission!

Stimulated transitions: the probability of them happening depends on the number of photons around

How did it all begin?

Rayleigh-Jeans law (circa 1900):

$$\text{energy density of a radiation field } u(\nu) = 8\pi\nu^2 kT/c^3$$

Note: the units of this expression are correct. Strictly speaking, $u(\nu)$ is an energy density per unit bandwidth, such that the integral $\int u(\nu) d\nu$ gives an answer with units of energy/volume.

Total energy radiated from a black body: $\int u(\nu) d\nu = \infty$

uh-oh... the "ultraviolet catastrophe"

Solution: quantum mechanics

Time-dependent perturbation theory

As a result of a perturbation $h(t)$, a system in quantum state 1 makes a transition to quantum state 2 with probability given by:

$$P_{1 \rightarrow 2} = \frac{1}{\hbar^2} \left| \int_{-\infty}^t e^{iE_{21}t'/\hbar} h(t') dt' \right|^2$$

Notation:

$$\omega_{21} = \frac{E_{21}}{\hbar}$$

Harmonic perturbation

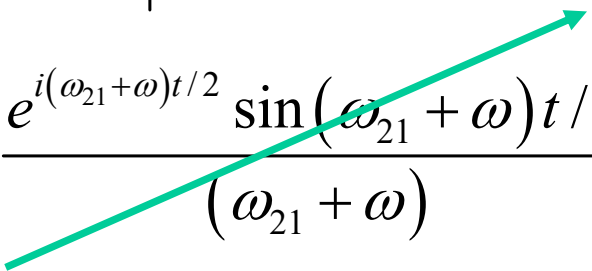
Key example: suppose we subject a two-level system, initially in state 1, to a harmonic perturbation, of the form:

$$h(t) = \begin{cases} 0 & t < 0 \\ 2A_0 \sin \omega t & t > 0 \end{cases} \quad \text{(and suppose that the frequency of the perturbation, } \omega, \text{ is close to } \omega_{21})$$

Transition probability to state 2 is:

$$P_{1 \rightarrow 2} = \frac{A_0^2}{\hbar^2} \left| \int_0^t e^{i\omega_{21}t'} (e^{i\omega t'} - e^{-i\omega t'}) dt' \right|^2$$

$$= \frac{4A_0^2}{\hbar^2} \left| \left(\frac{e^{i(\omega_{21}-\omega)t/2} \sin(\omega_{21}-\omega)t/2}{(\omega_{21}-\omega)} \right) - \left(\frac{e^{i(\omega_{21}+\omega)t/2} \sin(\omega_{21}+\omega)t/2}{(\omega_{21}+\omega)} \right) \right|^2$$

RWA 

$$\approx \frac{4A_0^2}{\hbar^2} \cdot \frac{\sin^2 [(\omega_{21}-\omega)t/2]}{(\omega_{21}-\omega)^2}$$

Note that $P_{1 \rightarrow 2} = P_{2 \rightarrow 1}$

Absorption and stimulated emission are **equally likely!**

Einstein A and B coefficients

Consider a radiation field and a collection of two-level systems, in thermal equilibrium with each other.

stimulated emission probability: proportional to the number of atoms in upper state N_2 , and also to the number of photons

spontaneous emission probability: proportional to N_2 , but does not depend on the photon density!

$$W_{2 \rightarrow 1} = A \cdot N_2 + B \cdot u(\nu) N_2$$

Note: this is the same as γ_{rad}

stimulated absorption probability: proportional to the number of atoms in lower state N_1 , and also to the number of photons

spontaneous absorption: there is no such thing

$$W_{1 \rightarrow 2} = B \cdot u(\nu) N_1$$

Quantum mechanics says that these two coefficients must be equal!

But: in thermal equilibrium, the upward and downward transition rates must balance:

$$W_{1 \rightarrow 2} = W_{2 \rightarrow 1}$$

Einstein A and B coefficients

Equate these two rates:

$$\frac{N_1}{N_2} = \frac{A + B \cdot u(\nu)}{B \cdot u(\nu)}$$

But Boltzmann's Law tells us that
(in equilibrium)

$$\frac{N_1}{N_2} = e^{(E_2 - E_1)/kT}$$

Recognizing that $E_2 - E_1 = h\nu$, we solve for $u(\nu)$:

$$u(\nu) = \frac{A/B}{e^{h\nu/kT} - 1}$$

This must correspond to the Rayleigh-Jeans result in the classical limit ($h \rightarrow 0$), which implies:

$$\frac{A}{B} = \frac{8\pi h\nu^3}{c^3}$$

Also has units of energy density per unit bandwidth

Since $A = \gamma_{\text{rad}}$, we can now solve for B also:

$$B = \frac{\gamma_{\text{rad}} c^3}{8\pi h\nu^3}$$

Transition rates

Our expression for the downward transition rate is now:

$$\begin{aligned}W_{2 \rightarrow 1} &= A \cdot N_2 + B \cdot u(\nu) N_2 \\ &= A \cdot N_2 \left(1 + \frac{B}{A} u(\nu) \right)\end{aligned}$$

Bose-Einstein
distribution



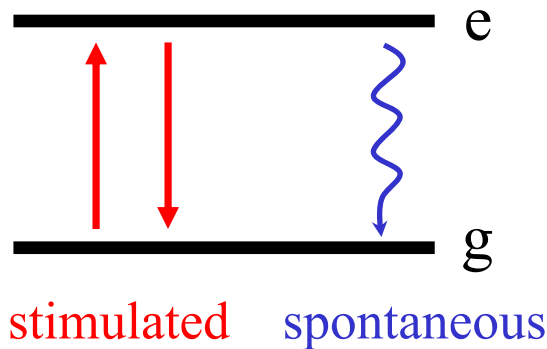
But since $u(\nu) = \frac{A/B}{e^{h\nu/kT} - 1}$ we therefore have $W_{2 \rightarrow 1} = A \cdot N_2 \left(1 + \frac{1}{e^{h\nu/kT} - 1} \right)$

In other words, W_{21} is proportional to: 1 + the number of photons.

It is easy to see that the upward transition rate, W_{12} , is proportional to the number of photons:

$$W_{1 \rightarrow 2} = B \cdot u(\nu) N_1 = AN_1 \left(\frac{1}{e^{h\nu/kT} - 1} \right)$$

Rate equation analysis



spontaneous emission:
proportional to initial state population

stimulated transitions:
proportional to initial state population
proportional to photon density n_p
the same for upwards, downwards transitions

$$\frac{dN_e}{dt} = -\gamma_{eg}N_e - Kn_pN_e + Kn_pN_g = -\frac{dN_g}{dt}$$

$$\frac{dN_e}{dt} = -\gamma_{eg}N_e + Kn_p(N_g - N_e)$$

Note: the constant K is simply given by $h\nu \cdot B$, where B is the Einstein B coefficient

Rate equation analysis, continued

$$\frac{dN_e}{dt} = -\gamma_{eg} N_e + Kn_p (N_g - N_e)$$

emitted photons go
in all directions

emitted photons go only into the
direction of the incident light

So, photon number varies according to:

$$\frac{dn_p}{dt} = -Kn_p (N_g - N_e) = -Kn_p \cdot \Delta N$$

→ Number of photons grows **exponentially** if $N_e > N_g$

A LASER!

This condition is known
as “population inversion”.

Rate equation analysis, part 3

In thermal equilibrium:

$$N_e = e^{-\Delta E/kT} \cdot N_g < N_g$$

→ Population inversion is impossible in equilibrium.

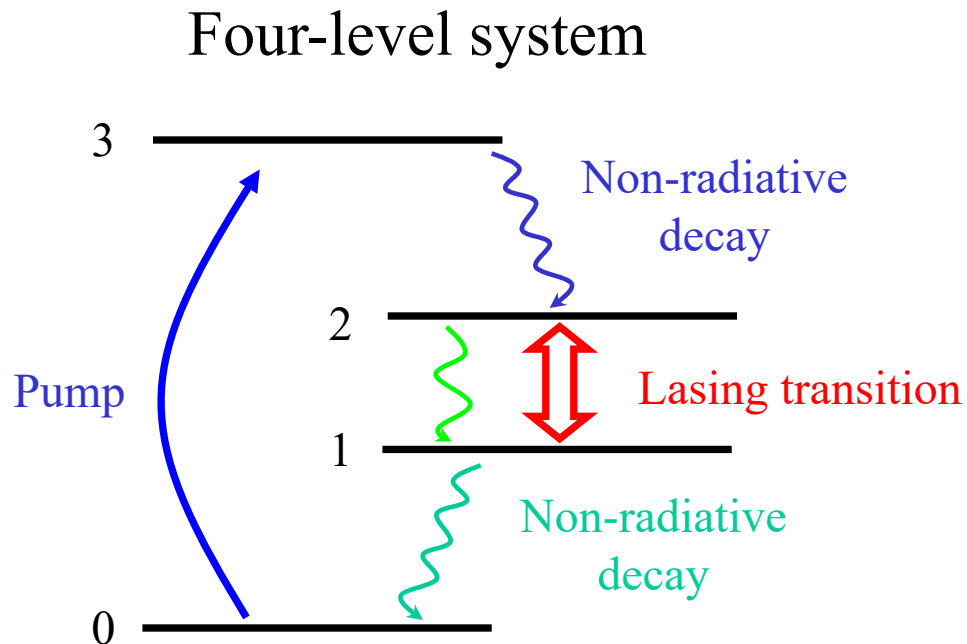
In a steady-state situation:

$$\frac{dN_e}{dt} = -\gamma_{eg}N_e + Kn_P(N_g - N_e) = 0$$

$$N_e = \frac{Kn_P}{\gamma_{eg} + Kn_P} N_g < N_g$$

→ Population inversion is impossible in steady-state.

So how do you make a laser?



Steps 1 and 2:

Combine to give an *effective* pumping rate for level 2: R_p

Step 3:

stimulated transitions due to n_p
spontaneous decay rate: γ_{21}

Step 4:

spontaneous decay rate: γ_{10}

$$\frac{dN_2}{dt} = R_p - \gamma_{21}N_2 + Kn_p(N_1 - N_2)$$

$$\frac{dN_1}{dt} = \gamma_{21}N_2 - Kn_p(N_1 - N_2) - \gamma_{10}N_1$$

$$\frac{dN_0}{dt} = \gamma_{10}N_1 - R_p$$

The four-level model

Steady-state solution:

$$\Delta N = N_1 - N_2 = \frac{R_P}{\gamma_{10}} \left[\frac{\gamma_{21} - \gamma_{10}}{\gamma_{21} + Kn_P} \right]$$

Population inversion (i.e., $\Delta N < 0$) is assured if $\gamma_{10} > \gamma_{21}$

(even if $n_p = 0$, and even if R_p is small)

A necessary condition for lasing

There are other necessary conditions, e.g.:

- a resonant cavity - provides feedback
- net gain per round trip $>$ net loss per round trip - “threshold”

Saturation in the four-level atom

$$\Delta N = \frac{R_P}{\gamma_{10}} \left[\frac{\gamma_{21} - \gamma_{10}}{\gamma_{21} + Kn_P} \right] = \frac{\gamma_{21} - \gamma_{10}}{\gamma_{21}\gamma_{10}} \cdot R_P \left[\frac{1}{1 + W_{sig} \tau_{21}} \right]$$

- population inversion when $\tau_2 > \tau_1$
- small signal inversion is proportional to the pump rate
- inversion level drops when $W_{sig} > \gamma_{21}$
- the characteristic intensity for this effect is independent of pump rate R_p

↑
 ΔN_0

↑ ↑
 Kn_p $1/\gamma_{21}$

→ "gain saturation"

Note: W_{sig} is proportional to n_p and therefore to the intensity of the light in the medium. Thus, we can define a **saturation intensity** I_{sat} such that:

$$W_{sig} \tau_{21} = I/I_{sat}$$

